

# Optimal & fast Wiener filtering of CMB maps without preconditioning



Doogesh Kodi Ramanah<sup>1,2</sup>, Guilhem Lavaux<sup>1,2</sup>, Benjamin D. Wandelt<sup>1,2</sup>

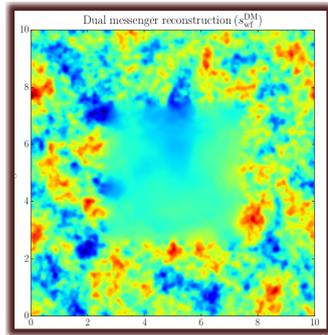
<sup>1</sup>IAP/CNRS, <sup>2</sup>ILP



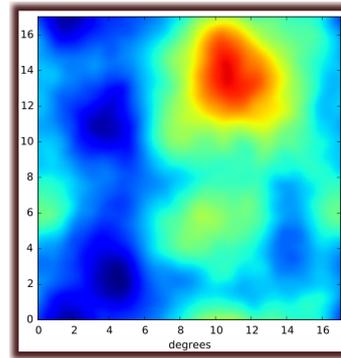
(DKR, Lavaux & Wandelt 2017)

(Anderes, Lavaux & Wandelt 2015)

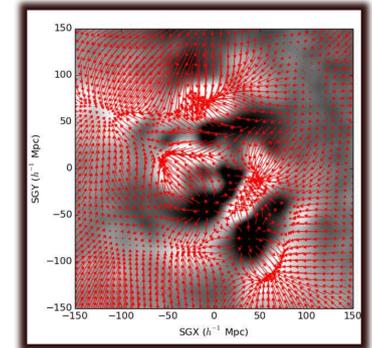
(Lavaux 2017, in prep.)



Map-making

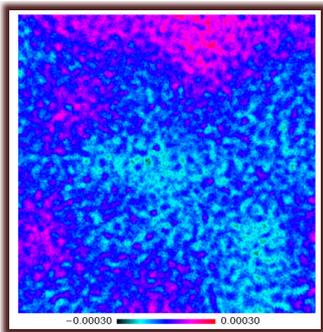


Lensing



Velocity field

(Elsner & Wandelt 2008)

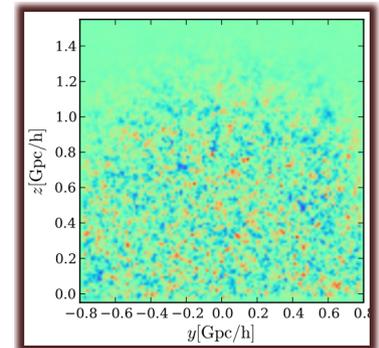


Non-Gaussianity

Wiener Filter

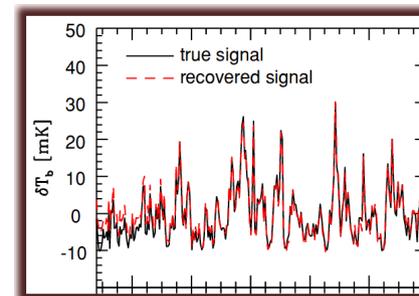
LSS

Density field



(Jasche & Lavaux 2015)

21-cm field



(Gleser+ 2008)

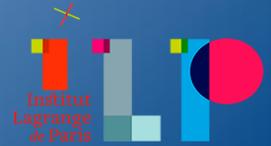
- Power spectrum estimation
- Likelihood analysis
- De-lensing
- Foreground removal
- Template-matching

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## WIENER FILTER

- Signal reconstruction from noisy data
- **Optimal** data analysis solution for Gaussian Random Fields  $\approx$  Cosmic density field

Wiener filter equation

$$(\mathbf{S}^{-1} + \mathbf{N}^{-1}) \mathbf{S}_{\text{wf}} = \mathbf{N}^{-1} \mathbf{d}$$

↑                      ↑                      ↑                      ↑  
Signal covariance    Noise covariance    Wiener filter solution    data

## Conventional Methods

- Preconditioned Conjugate Gradient (**PCG**)
- Preconditioner,
  - ✗ **Problem-dependent**
  - ✗ **Ill-conditioning issue**
  - ✗ **Numerically expensive & unstable**
- High level of complexity
- Difficulties in dealing with CMB polarization

## ★ DUAL MESSENGER ALGORITHM

- Fast & efficient preconditioner-free algorithm
- Deals effectively with CMB polarization data, while PCG fails to converge to a sensible solution
- Unconditionally stable & trivial numerical implementation
- Relevant for current & next-generation CMB experiments.

➡ Algorithm can be augmented to deal with more complex & realistic noise models

➡ Exact global Bayesian analyses (Gibbs sampling) for optimal separation of pure  $E$ - &  $B$ -modes

(DKR, Lavaux & Wandelt 2017, MNRAS) → [arXiv:1702.08852](https://arxiv.org/abs/1702.08852)